

A Search for
 $\nu_\mu \rightsquigarrow \nu_e$ Oscillations
in the NOMAD Experiment

Andrew R. Godley



*A thesis
submitted for the degree of
Doctor of Philosophy
at the
University of Sydney*

August, 2000

Abstract

The NOMAD experiment is a neutrino oscillation experiment, capable of identifying ν_μ , $\bar{\nu}_\mu$, ν_e , $\bar{\nu}_e$ and ν_τ for use in oscillation analyses. A search for $\nu_\mu \rightsquigarrow \nu_e$ oscillations is conducted, emphasising the development of two separate beam simulators, to provide the background, (no oscillation), ν_e signal.

Both beam descriptions include fits to the results of the SPY experiment that measured hadron production from a 450 GeV proton beam on beryllium target. An independent analysis of the raw SPY data to produce the particle yield is reported.

A series of criteria are described for the selection and classification of neutrino events. These produce the data samples necessary for both tuning the beam simulation and determining the oscillation signal.

The development of a GEANT and FLUKA based Monte Carlo beam simulator is presented, providing good agreement to the measured neutrino beam. This simulation method has sizeable variations depending on the beamline geometry, which is not known precisely. This causes large systematic errors.

An empirical parametrisation is proposed and used for the first time in a NOMAD oscillation search. It uses the measured neutrino spectra at NOMAD, except the ν_e , to infer the meson production at the target, and then predict the ν_e spectrum. This method has good agreement with the data and is also insensitive to alterations of the beamline geometry, resulting in much smaller systematic errors.

The reduction of the systematic errors allows the $\nu_\mu \rightsquigarrow \nu_e$ oscillations search to be performed with much greater precision. Comparisons of the ν_e/ν_μ ratio between the empirical parametrisation and data yields no evidence for $\nu_\mu \rightsquigarrow \nu_e$ oscillations, setting a limit on the mixing parameter, $\sin^2(2\theta) < 1.9 \times 10^{-3}$ (90% CL) at high Δm^2 . The present sensitivity of the analysis on the mixing parameter is 0.0017.

Acknowledgements

I would like to say a few words to thank those involved in helping me with this thesis. Firstly there is my supervisor, A/Prof Lawrence Peak who warmly brought me into Falkiner High Energy Physics, and since then has been equally thoughtful and helpful in his supervision of this work and my schedule. Dr Kevin Varvell must also be gratefully thanked as his guidance was invaluable when I started this project and he continued answering my queries endlessly. The other members of Falkiner kept the work place pleasant and amiable, particularly during outings and talks. They are Drs Andy Bakich, Grant Gorfine, Reza Hashemi and Juris Ulrichs, (with whom I shared a flat in St Genis.)

There are many people at CERN who I would also like to thank. Dr Emmanuel Tsesmelis took the role of my supervisor during my first excursion to CERN in 1997. I owe a great deal to him for supporting my work in SPY and the neutrino beam studies, whilst also for presenting my work at CERN when I was in Sydney. When my focus shifted to the empirical parametrisation and the oscillation search I received great help from Dr Antonella de Santo, with whom I worked on the event selection and the empirical fitting programmes. Dr Fred Weber was also involved in these projects. Dr Sanjib Mishra provided deeper insights into the oscillation search, its requirements, goals and pitfalls. I am grateful to him for his help. Of course I was greatly helped by Dr Paul Soler who frequented both CERN and Falkiner in my time. Paul took interest in my projects, presented my work, reliably kept me up to date with the happenings at CERN and provided any extra aid I needed.

I had the pleasure of spending time with many friendly students. At Falkiner there was Bruce Canu, Malcolm Ellis, Aldo Saavedra, Joe Seisdodos, Jaime Varas and Jianguai Wang. Bruce provided all the distractions that one needed. Malcolm and Joe were officemates, available for conversation on all topics. Aldo lightened up my whole time in Sydney, both at work and after, whilst Jaime regaled everyone with stories of his exploits. I would also like to thank the two former students, Drs Steven Boyd and Bruce Yabsley, both of whom helped me a great deal. Amongst many other things, Steve provided expertise when I was studying the FCAL and Bruce advice on L^AT_EX. At CERN I would like to thank Jukka Kokkonen and Dr Kai Zuber. From Melbourne University I would like to thank Brendan Dick, for many a relaxing sunday afternoon whilst at CERN. A special hello goes to my fellow Board Members at St Genis Uni.

To everyone mentioned, and those I forgot, thank you very much!

Author's Contribution

This work includes contributions from other members of the NOMAD collaboration. These contributions are acknowledged below on a chapter by chapter basis. Also note that the introduction is merely the author's summary of readily available neutrino theory and experimental results, given to provide a background for the project.

The description of the NOMAD experiment arises from the author's working knowledge of the systems, as well as published papers and internal papers, duly acknowledged. This is also true for the description of the SPY experimental setup.

The study of the SPY data is the author's work. The analysis of G.M. Collazuol and A. Guglielmi was used as a springboard, and was subsequently ratified during the author's investigation.

The neutrino event classification comprises work conducted by the author and A. DeSanto. The Calisto selection programme was written and maintained by the author, whilst Readall was developed by A. DeSanto until December 1999, and then by the author. The selection criteria are based on the studies of the NOMAD Phase II group, further refined by A. DeSanto, S.R. Mishra and the author.

The testing of NUBEAM involved a great deal of the author's time. Whilst most developments were implemented into the NUBEAM code by D. Daniels, F.J.P. Soler, V. Valuev and T. Weiße, tests and comparisons were carried out by the author, amongst others. A. Ferrari was responsible for the development of FLUKA, and G.M. Collazuol and A. Guglielmi for the SPY weighting, and again testing was undertaken by the author. The comparison of simulations to the muon pit data was conducted by the author alone, within NOMAD. The FCAL analysis was entirely the author's work, along with the comparison of NUBEAM to the drift chamber data. The systematic error study reported is that of V. Valuev.

The concept of the empirical parametrisation in NOMAD is S.R. Mishra's. The description given is of the author's implementation of the concept, to obtain a parametrisation independently to S.R. Mishra. The initial fitting programme was set up by F. Weber. The author made the extensions and adaptations necessary to include the simplex fitting mechanism, $\bar{\nu}_\mu$ and $\bar{\nu}_e$ fitting, the fit to SPY data and the production of the ν_e estimate. The independent prediction of the ν_e flux formed a major part of the work for this thesis.

The last chapter is the author's analysis, with the inclusion of S.R. Mishra's beam prediction and small contributions noted therein.

Preface

This thesis is a report on work carried out for the degree of Doctor of Philosophy in the Falkiner High Energy Physics Department at the University of Sydney. The work was undertaken as part of the NOMAD Collaboration at CERN. This thesis describes the NOMAD experiment and the analysis of its data for indications of muon neutrino to electron neutrino oscillations, featuring the prediction of the neutrino beam by the empirical parametrisation, in a blind analysis. Additional studies, relevant to the oscillation search, are also presented.

The first chapter introduces the theory of neutrinos. A history is given as to their importance in particle physics and cosmology. The concept and relevance of a massive neutrino is introduced, and neutrino oscillations are examined as an avenue for determining neutrino mass.

Chapter 2 describes the NOMAD experiment. The neutrino oscillation search technique is summarised, which indicates the necessary requirements of NOMAD. The neutrino beam is briefly discussed, followed by a rundown on the subdetectors that comprise NOMAD. The NOMAD triggering, slow control and data taking are detailed, along with a description of event reconstruction and simulation in NOMAD.

The stand-alone experiment, SPY, is described in chapter 3. This experiment measured meson production relevant to the NOMAD neutrino beam. The layout of the SPY experiment is described, and then an analysis of its data is presented, including particle identification and yield computations. The results of this analysis and an indication of the future uses of the SPY data in neutrino beam simulation and parametrisation are given.

The fourth chapter deals with the classification of neutrino interactions within NOMAD. It presents the selection criteria for accepting and classifying neutrino triggers and also tabulates the efficiency of the selection process, which is later used in data to simulation comparisons. The event samples produced by these criteria are used throughout this thesis. The determination of the number of electron and muon neutrinos in this chapter, along with the beam predictions of the two following chapters, form the two parts of the oscillation search.

Chapter 5 summarises the traditional effort to simulate the neutrino beam with the programmes NUBEAM and FLUKA, and with fits to the SPY data. The chapter explains why the neutrino beam is so critical to the oscillation search and the available means for testing its simulation, including the muon pit data and quasielastic-like events in the FCAL. The three beam focusings used are described, and then the results of their simulation are compared with the NOMAD data.

The empirical parametrisation is proposed in chapter 5 as a new and novel beam flux prediction method. The theory, practice and merits of the method are described. A parametrisation is then fully determined from fits to SPY and NOMAD data. This parametrisation is cross checked with the parametrisation of S.R. Mishra in the next chapter.

The work of chapters 4, 5 and 6 are combined in chapter 7 to determine the oscillation signal. A study of systematic errors and a discussion of why the empirical parametrisation is better suited to the oscillation search than NUBEAM alone are included. The principles of a blind analysis are highlighted, before the oscillation signal is searched for by the comparison of the expected ν_e/ν_μ energy spectrum with that from data. The results are then quoted and discussed, before being reiterated in the conclusion.

Contents

Abstract	i
Acknowledgements	ii
Author's contribution	iii
Preface	iv
Contents	vi
List of Figures	xi
List of Tables	xvi
1 Introduction and Theory	1
1.1 Introduction	1
1.2 Neutrinos in Cosmology	3
1.2.1 Relic Neutrinos	3
1.2.2 Dark Matter	3
1.3 Neutrino Physics	5
1.3.1 Dirac and Majorana Neutrinos	5
1.3.2 Neutrino Interactions	6
1.3.3 Neutrino Sources	8
1.4 Neutrino Oscillations	10
1.5 The Neutrino Mass Situation	12
1.5.1 The See-Saw Mechanism	12
1.5.2 Direct Mass Measurements	13
1.5.3 Oscillation Search Results	14
1.5.4 The Ongoing Search	18
2 The NOMAD Experiment	21
2.1 Introduction	21
2.2 Oscillation Searches	22
2.2.1 Muon Neutrino to Tau Neutrino	22
2.2.2 Muon Neutrino to Electron Neutrino	23
2.3 NOMAD as a check on LSND	23
2.4 The Neutrino Beam	24
2.4.1 Proton Acceleration and Beam Cycle	24

2.4.2	Beam Monitoring	26
2.4.3	Target	27
2.4.4	Neutrino Beam Generation	27
2.4.5	Focusing and Decay Tunnel	28
2.4.6	Muon Pits and Shielding	29
2.4.7	Expected Beam	29
2.5	Detector Description	29
2.5.1	The Magnet	30
2.5.2	Veto Counters	32
2.5.3	Forward Calorimeter	33
2.5.4	Drift Chamber	34
2.5.5	Transition Radiation Detector	36
2.5.6	Trigger Planes	38
2.5.7	Preshower	39
2.5.8	Electromagnetic Calorimeter	40
2.5.9	Hadronic Calorimeter	41
2.5.10	Muon Chambers	42
2.6	NOMAD Control Systems	44
2.6.1	Data Acquisition	44
2.6.2	Trigger Conditions	45
2.6.3	Slow Controls	47
2.7	Simulation and Reconstruction of Events	47
2.7.1	Event Generation	48
2.7.2	Detector Simulation	48
2.7.3	Event Reconstruction	49
2.8	The NOMAD Data Set	49
3	The SPY Experiment	51
3.1	Introduction and Motivation	51
3.2	Experimental Layout	52
3.2.1	Beam and Target	52
3.2.2	The H6 Spectrometer	54
3.2.3	Data Acquisition and Triggering	58
3.3	Particle Identification	59
3.3.1	Electron and Muon Identification	59
3.3.2	Intermediate Momenta	60
3.3.3	High Momenta	62
3.4	Production of Yields	64
3.4.1	Detector and Selection Efficiencies	64
3.4.2	Prescales	64
3.4.3	Acceptance	66

3.4.4	Correction for Particle Decays	67
3.4.5	Strange Particle Decays	68
3.4.6	Normalisation to Protons on Target	69
3.4.7	Empty Target Subtraction	69
3.5	Results and Comparisons	71
3.5.1	Momentum Dependence	72
3.5.2	Angular Dependence	73
3.5.3	Target Length Dependence	73
3.5.4	Other SPY Results	79
4	Neutrino Event Classification	81
4.1	Common Lepton Selection Criteria	82
4.1.1	Quality and Filter Cuts	82
4.1.2	Vetoing Through-Going Muons	85
4.1.3	The LEP Condition	85
4.1.4	Prompt Lepton Selection	86
4.2	Muon Identification	87
4.3	Electron Identification	89
4.3.1	Phase II Selection	89
4.3.2	Prompt Electron Selection	92
4.3.3	Increased Background Rejection	94
4.4	Selection Efficiencies	95
4.4.1	Energy Smearing	103
4.5	Hadronic Energy Scale	105
5	Simulation of the Neutrino Beam	111
5.1	Development of NUBEAM	112
5.1.1	Early Versions and Prehistory	113
5.1.2	Version 4.00	114
5.1.3	Version 5.00	114
5.1.4	Version 6.00	117
5.2	Development of FLUKA	120
5.2.1	FLUKA Versions	121
5.2.2	SPY Weighted FLUKA	123
5.3	Running NUBEAM and FLUKA	125
5.3.1	Interfacing FLUKA to NUBEAM	125
5.3.2	NUBEAM Decay Multiplicity	125
5.3.3	Weighting Event Monte Carlo	126
5.4	Muon Pit Comparisons	127
5.4.1	Layout of the Muon Pits	128
5.4.2	Calibration	129

5.4.3	Comparison with Simulation	130
5.5	Quasielastic Like Events in The FCAL	131
5.5.1	Further Details of the FCAL	133
5.5.2	Low ν Events	134
5.5.3	Event Selection and Efficiency	136
5.5.4	Comparison with Simulations	139
5.6	Drift Chamber Events	140
5.6.1	Positive, Negative and Zero Focusing Beams	140
5.6.2	Data Monte Carlo Comparisons	142
5.6.3	Comparison with CHORUS Data and Simulation	148
5.7	Systematic Errors	152
5.8	Summary of Improvements	155
6	Empirical Parametrisation of the Neutrino Beam	157
6.1	Introduction and Motivation	157
6.2	Theoretical Concept	158
6.3	Implementation	159
6.3.1	Data Sample	159
6.3.2	Target Monte Carlo Sample	160
6.3.3	Meson Production Parametrisation	163
6.3.4	Corrections to the Estimate	165
6.3.5	Normalisation of Pion and Kaon Spectra	169
6.3.6	Mechanism to Produce a Trial Flux Estimate	171
6.3.7	Simplex Minimisation	173
6.3.8	SPY Data Fit	174
6.3.9	Producing the ν_e Estimate	176
6.4	Results	180
6.4.1	ν_μ Fit	180
6.4.2	$\bar{\nu}_\mu$ Fit	185
6.4.3	$\bar{\nu}_e$ Fit	193
6.4.4	ν_e Prediction	193
7	Oscillation Search	199
7.1	Discussion of Beam Simulators	199
7.1.1	Empirical Parametrisations	200
7.1.2	Official Parametrisation Compared to Data	203
7.1.3	Official Parametrisation Compared to NUBEAM	205
7.1.4	Official Parametrisation Systematics	211
7.1.5	Benefits of the Empirical Parametrisation	215
7.2	Theory of Small Signal Searches	215
7.2.1	Blind Analysis and Consequences	215

7.2.2	Double Ratio and Analysis Binning	216
7.2.3	Statistical Approach	217
7.2.4	Oscillation Contribution	218
7.3	Statistical Errors	219
7.4	Systematic Errors	219
7.4.1	Event Classification	219
7.4.2	Background	221
7.4.3	Electron Selection	221
7.4.4	Energy Reconstruction	222
7.4.5	Summary of all Systematic Errors	223
7.5	Results	223
7.5.1	Measured Oscillation Signal	225
7.5.2	Significance	227
7.5.3	Future Work	228
	Conclusion	233
	Bibliography	235

List of Figures

1.1	The contribution of neutrino mass to the relative density . . .	5
1.2	Neutrino nucleon interaction Feynman diagrams	7
1.3	$\nu_\mu \rightsquigarrow \nu_e$ phase space	19
2.1	Layout of the WANF optimised for NOMAD	25
2.2	Cycle of the SPS	26
2.3	The effect of the horn and reflector in focusing particles of varying energy	28
2.4	Neutrino spectra at NOMAD	30
2.5	Detector layout (1995-96)	31
2.6	Layout of NOMAD veto planes	32
2.7	Top view of the FCAL	34
2.8	A single drift chamber, with detail of a detection gap	35
2.9	Schematic view of a TRD module	37
2.10	Trigger plane layout	38
2.11	An exploded view of the Preshower	39
2.12	Schematic view of the lead glass blocks that comprise the ECAL	41
2.13	The front view of the HCAL	42
2.14	Muon chamber layout	43
2.15	The standard configuration of the NOMAD data acquisition .	45
2.16	Accumulated BCT on the T9 target	50
3.1	The T4 target station in plan view	54
3.2	The SPY spectrometer, shown in the horizontal and vertical planes	56
3.3	Energy deposited in the first module, as a fraction of the total deposited, plotted against the total deposited	60
3.4	The mass squared fit produced by the time of flight counters at 20 GeV/c	61
3.5	Background from empty target runs	70
3.6	Background yields	72

3.7	Momentum distribution of particle yields from a 100 mm target, at 0 mrad	74
3.8	Momentum distribution of particle ratios from a 100 mm target, at 0 mrad	75
3.9	The angular distribution of particle yields from a 100 mm target at 40 GeV/c (left) and 15 GeV/c (right)	76
3.10	The angular distribution of particle yields from a 100 mm target at 40 GeV/c (left) and 15 GeV/c (right)	77
3.11	Particle yields from 100, 200 and 300 mm target lengths at 15 and 40 GeV/c, in the forward direction	78
4.1	A candidate ν_μ CC event as seen in the yz view of the NOMAD event display	88
4.2	Likelihood distributions for simulated 10 GeV pion and electron tracks crossing all nine TRD modules	90
4.3	Candidate ν_e CC event, as viewed in the NOMAD event display	95
4.4	Reconstruction differences in the z vertex position	96
4.5	The ν_μ DIS selection efficiency as a function of generated neutrino energy.	99
4.6	The $\bar{\nu}_\mu$ DIS selection efficiency as a function of generated neutrino energy.	100
4.7	The ν_e DIS selection efficiency as a function of generated neutrino energy.	102
4.8	The $\bar{\nu}_e$ DIS selection efficiency as a function of generated neutrino energy.	102
4.9	The ν_μ QE selection efficiency as a function of generated neutrino energy.	104
4.10	The ν_μ RES selection efficiency as a function of generated neutrino energy.	105
4.11	Scatter plot of E_{vis} versus E_ν for ν_μ DIS CC Monte Carlo	106
4.12	Determining the hadronic energy scale	107
4.13	Monte Carlo and data ν_μ spectra without the hadronic energy scale	109
4.14	Monte Carlo and data ν_μ spectra with the hadronic energy scale	110
5.1	Changes due to version 4.04 in the neutrino energy spectra for the four types	115
5.2	Version 5.00 changes in the neutrino energy spectra for the four types	116
5.3	The distribution of primary proton interaction points in the decay tunnel	118

5.4	The distribution of primary proton interaction points that lead to a ν_μ or $\bar{\nu}_\mu$ at NOMAD	119
5.5	Version 6.12 changes in the neutrino energy spectra for the four types	121
5.6	Positive focusing NUBEAM 6.12 weight functions	127
5.7	Layout of the first three muon pits	129
5.8	Simulation of the angular distribution of muons and electrons in the calibration emulsion of muon pit 3	131
5.9	Comparison of the data to simulated x and y spatial distributions of the muon flux in the pits	132
5.10	Side view of the FCAL	135
5.11	Selection efficiency of quasielastic (QE) neutrino events in the FCAL as a function of reconstructed neutrino energy.	138
5.12	Spatial distributions of low ν events in the FCAL	139
5.13	Energy spectrum of low ν events in the FCAL compared to Monte Carlo flux predictions of NUBEAM 6.12 with FLUKA 98.	140
5.14	Negative focusing NUBEAM 6.12 weight functions	142
5.15	Zero focusing NUBEAM 6.12 weight functions	143
5.16	Positive focusing ν_μ DIS drift chamber data and simulation	145
5.17	Positive focusing $\bar{\nu}_\mu$ DIS drift chamber data and simulation	146
5.18	Positive focusing $\bar{\nu}_e$ DIS drift chamber data and simulation	147
5.19	Negative focusing $\bar{\nu}_\mu$ DIS drift chamber data and simulation	149
5.20	Negative focusing ν_μ DIS drift chamber data and simulation	150
5.21	Zero focusing ν_μ DIS drift chamber data and simulation	151
5.22	Zero focusing $\bar{\nu}_\mu$ DIS drift chamber data and simulation	152
6.1	Neutrino spectra separated according to neutrino parent	159
6.2	The five radial bins and the drift chamber.	161
6.3	Selection efficiency for ν_μ as a function of true neutrino energy	166
6.4	ν_μ spectrum showing the approximate energy ranges where pions and kaons dominate	170
6.5	Fits to SPY K^+/π^+ data	177
6.6	Fits to SPY K^-/π^- data	178
6.7	NOMAD ν_μ data versus the empirical parametrisation in the five radial bins, linear	182
6.8	NOMAD ν_μ data versus the empirical parametrisation in the five radial bins, logarithmic	183
6.9	NOMAD ν_μ data versus empirical parametrisation in the combined radial bin	184

6.10	NOMAD ν_μ negative focusing data versus empirical parametrisation in the combined radial bin	186
6.11	NOMAD $\bar{\nu}_\mu$ data versus the empirical parametrisation in the five radial bins, linear	187
6.12	NOMAD $\bar{\nu}_\mu$ data versus the empirical parametrisation in the five radial bins, logarithmic	188
6.13	NOMAD $\bar{\nu}_\mu$ data versus empirical parametrisation in the combined radial bin	189
6.14	NOMAD $\bar{\nu}_\mu$ negative focusing data versus the empirical parametrisation in the five radial bins, linear	190
6.15	NOMAD $\bar{\nu}_\mu$ negative focusing data versus the empirical parametrisation in the five radial bins, logarithmic	191
6.16	NOMAD $\bar{\nu}_\mu$ negative focusing data versus empirical parametrisation in the combined radial bin	192
6.17	NOMAD $\bar{\nu}_e$ data versus empirical parametrisation in the combined radial bin, linear	194
6.18	NOMAD $\bar{\nu}_e$ data versus empirical parametrisation in the combined radial bin, logarithmic	195
6.19	The empirical parametrisation prediction of ν_e , linear	196
6.20	The empirical parametrisation prediction of ν_e , logarithmic	197
7.1	Independent and official parametrisation predictions of ν_μ CC low ν events	201
7.2	Ratio of independent over official parametrisation prediction of the ν_e/ν_μ ratio at low ν	202
7.3	The ν_μ energy spectra (high ν) of data and the official parametrisation	204
7.4	The $\bar{\nu}_\mu$ energy spectra (high ν) of data and the official parametrisation	206
7.5	The $\bar{\nu}_e$ energy spectra (high ν) of data and the official parametrisation	207
7.6	Official parametrisation and NUBEAM predictions of ν_μ CC high ν events	208
7.7	Official parametrisation and NUBEAM predictions of $\bar{\nu}_\mu$ CC high ν events	209
7.8	Official parametrisation and NUBEAM predictions of $\bar{\nu}_e$ CC high ν events	210
7.9	Official parametrisation and NUBEAM predictions of ν_e CC high ν events	212
7.10	Comparison of the ν_e/ν_μ ratio as predicted by NUBEAM and the official parametrisation	213

7.11	$R_{e\mu}$ determined with the standard energy scale and a 10% lower scale	223
7.12	$R_{e\mu}$ from the 1995 high ν data (points) and as predicted by the official parametrisation (histogram)	224
7.13	$R_{e\mu}$ from the 1995 low ν data (points) and as predicted by the official parametrisation (histogram)	225
7.14	The 90% confidence limit obtained for $\nu_\mu \rightsquigarrow \nu_e$ oscillations from the 1995 NOMAD data	226
7.15	The exclusion region of the analysis and the original LSND allowed region	227
7.16	The exclusion region of the analysis and the revised LSND allowed region	229
7.17	The 90% confidence level sensitivity expected from this analysis, using the entire NOMAD data set.	230
7.18	The revised allowed region of LSND compared to the sensitivity expected from this analysis, including the entire NOMAD data set.	231

List of Tables

1.1	Ranges of Δm^2 applicable to the four neutrino sources.	12
2.1	Summary of Data taken by NOMAD from 1995 to 1998	50
3.1	Summary of the different SPY running conditions	53
3.2	Particles identified in each of the four timing subgroups for a 20 GeV/c run.	62
3.3	Particles identified in each of the six timing subgroups for a 40 GeV/c run.	64
3.4	Summary of identified particles	65
3.5	Definitions of the trigger scalers used in SPY	66
3.6	The calculated spectrometer acceptances used in this analysis	68
3.7	Summary of identified particles in the empty target runs	71
4.1	Minimum z allowed for the fiducial volume.	84
4.2	Maximum values of hits for the through-going muon veto, ac- cording to the number of tracks in the event.	86
4.3	Maximum allowable scattering angles in radians for leptons, according to the number of tracks in the event.	87
4.4	Monte Carlo ν_e DIS events failing the standard selection which then either fail the LEP adjusted selection or are recovered.	94
4.5	Effect of the Calisto event filter on the data	97
4.6	Selection efficiencies for ν_μ and $\bar{\nu}_\mu$ DIS events	98
4.7	Selection efficiencies for ν_e DIS events	101
4.8	Selection efficiencies for $\bar{\nu}_e$ DIS events	103
4.9	Efficiency for the selection of events from RES and QE Monte Carlo samples	104
5.1	he changes to the neutrino spectra shape and size with NUBEAM version	116
5.2	Effect of FLUKA upgrades on the neutrino beam at NOMAD	123
5.3	Data and simulated muon fluxes	133

5.4	FCAL quasielastic event selection	137
5.5	Expected ratios of charged current neutrino events expected in the beam	144
5.6	The average DIS neutrino energy and number of events for positive focusing data and simulation	148
5.7	The average DIS neutrino energy and number of events for negative focusing data and simulation	148
5.8	The average DIS neutrino energy and number of events for zero focusing data and simulation	148
5.9	The predictions of GBEAM 99 versus NUBEAM 5.00	153
5.10	NOMAD and CHORUS data comparison	153
5.11	Ratios of ν_e/ν_μ in the modified field versions to the standard version	155
5.12	Ratios of ν_e/ν_μ in the modified material versions to the stan- dard version	156
5.13	Ratios of ν_e/ν_μ in the modified alignment versions to the stan- dard version	156
6.1	Results of event selection, with and without the $\nu < 5$ GeV condition.	160
6.2	Normalisation factors for combining the resonance, quasielas- tic and deep inelastic Monte Carlo samples.	162
6.3	Base parameters entered in NUBEAM for meson production.	165
6.4	Scales to the backgrounds to the $\bar{\nu}_\mu$, ν_e , and $\bar{\nu}_e$ signals for different ranges of ν	169
6.5	Pion and kaon normalisation factors, for each of the five radial bins and combined bin (cm).	171
6.6	SPY K/π production, versus those obtained from the fit, dis- played according to secondary momentum.	176
6.7	SPY K^+/π^+ production, versus those obtained from the fit, displayed according to p_T at the two given secondary momenta.	179
6.8	SPY K^-/π^- production, versus those obtained from the fit, displayed according to p_T at the two given secondary momenta.	179
6.9	Final meson production parameters after completion of the fitting process.	181
6.10	χ^2 of the fitted flux to the ν_μ data	181
6.11	χ^2 of the fitted flux to the $\bar{\nu}_\mu$ data	185
7.1	The breakdown of the neutrinos in terms of their parents	203
7.2	Comparison of the number of high ν CC events in data with the predictions of NUBEAM and the official parametrisation.	211

7.3	Energy binning for $R_{e\mu}$, (GeV).	217
7.4	Percentage statistical errors for the data and Monte Carlo samples used in the oscillation search	219
7.5	Cuts and their variations used in the systematic error studies.	220
7.6	Event numbers and efficiencies used to calculate the unfolded number of ν_μ and ν_e events.	220
7.7	The unfolded number and ratio of ν_μ and ν_e events	221
7.8	The contributions to the systematic error on $R_{e\mu}$	224